


Astronomy, Postage Stamps, Conics, and Eccentricity

Unit Analysis

Transparency 1 shows the Cone Nebula as it appeared in USA Today. We can discuss how technology has improved what we see in the universe, and that people are debating today what it is that we are seeing. This overhead uses the Cone Nebula as a lead-in to unit analysis, or scientific notation.

Unit Analysis

The Cone Nebula as photographed by the Hubble telescope in April 2002.



The distance across the photograph is about 2.5 light-years.

A light-year is the distance light travels in a year.

If we assume light travels 186,000 miles in one second, we can find the number of miles in 1 light-year by converting 186,000 miles/second to miles/year.

$$\frac{186,000 \text{ miles}}{1 \text{ sec}} \cdot \frac{60 \text{ sec}}{1 \text{ min}} \cdot \frac{60 \text{ min}}{1 \text{ hr}} \cdot \frac{24 \text{ hr}}{1 \text{ day}} \cdot \frac{365 \text{ days}}{1 \text{ year}}$$

$$= 5,865,696,000,000 \text{ miles/year}$$

Transparency 1

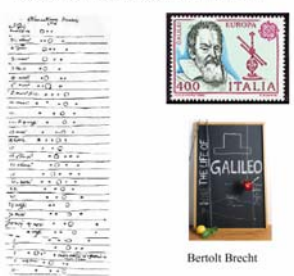
Scientific Notation

Put up Transparency 1 that was used previously, and remind them that the photograph was taken with the Hubble telescope. As technology has improved, so has our ability to see more of the details in space. What was the first technology that allowed us to improve our view of the sky? Put up Transparency 2. The table is from Galileo's notebook and shows the positions of four of Jupiter's moons. This leads into Transparency 3, which is a group project that will require students to work with numbers in scientific notation. Show the website that demonstrates Kepler's third law.

We cannot teach people anything; we can only help them discover it within themselves
-Galileo Galilei

Built the first telescope in 1609

Discovered the moons of Jupiter in 1610.





Bertolt Brecht

Transparency 2

Group Project

Using Jupiter's moons
to discover Kepler's third law.

Jupiter's Moon	Period T (seconds)	Radius R (km)	T^2/R^3
Io	1.53×10^5	4.22×10^3	
Europa	3.07×10^5	6.71×10^3	
Ganymede	6.18×10^5	1.07×10^4	
Callisto	1.44×10^6	1.88×10^4	

Transparency 3

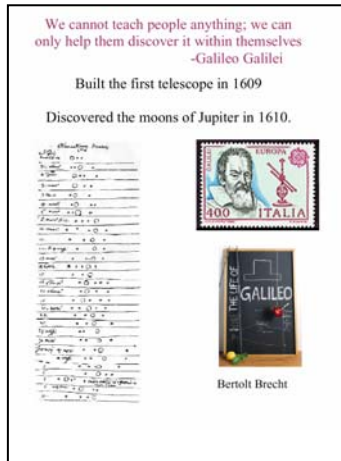
Merriam-Webster's COLLEGIATE DICTIONARY

One entry found for **scientific notation**.
Main Entry: **scientific notation**
Function: *noun*
: a widely used floating-point system in which numbers are expressed as products consisting of a number between 1 and 10 multiplied by an appropriate power of 10

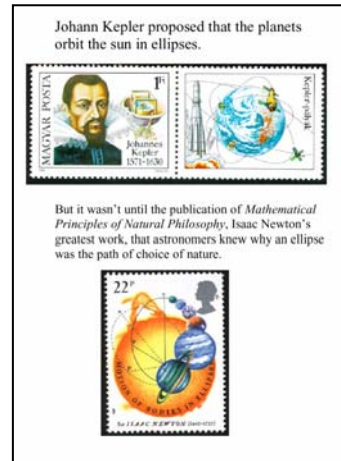
Transparency 21

Conic Sections

We introduce the conic sections by starting with just the ellipse. To start, put up Transparency 2 again to remind them that the first telescope was built in 1609. This leads into Transparency 4 which shows that in 1609 Kepler proposed that the planets orbit the sun in elliptical orbits. Kepler used empirical data for his conjecture. It took Newton to actually prove that a planetary orbit had to be a conic section.



Transparency 2



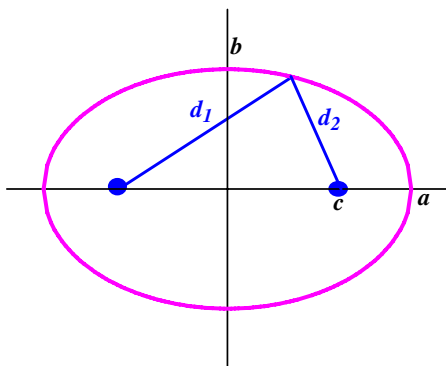
Transparency 4

Next, show the website demonstrating Kepler's first and second laws, so they can see an ellipse. Then draw an ellipse using the snaps and string on the overhead projector. To remind you of how we did that in the talk, here are the steps:

Deriving the Rectangular Equation for an Ellipse There are no transparencies for this page. This is just to remind you of the way we developed the initial facts about an ellipse using the string and snaps to draw an ellipse.

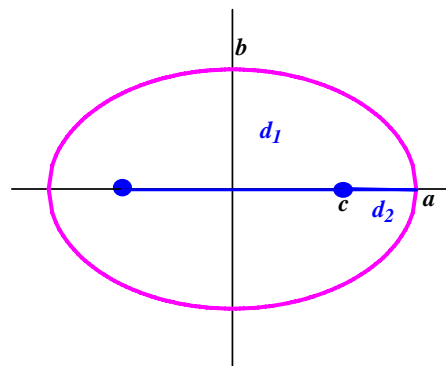
Step 1 After drawing an ellipse with the string, draw d_1 and d_2 out to an arbitrary point and write

$$d_1 + d_2 = \text{constant}$$



Step 2 Then move the arbitrary point to a to show that

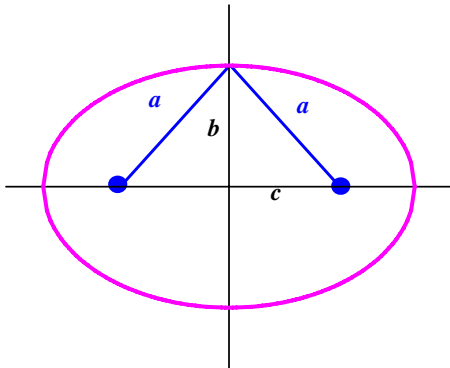
$$d_1 + d_2 = 2a$$



Conic Sections (Continued)

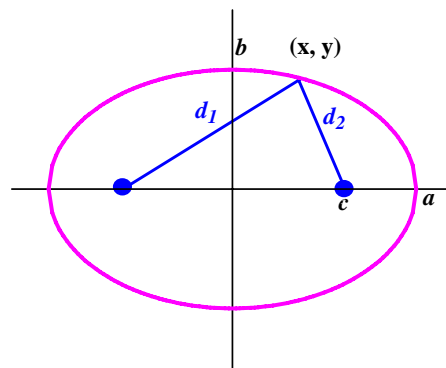
Step 3 Next, move the arbitrary point to the top of the ellipse to show that

$$a^2 = b^2 + c^2$$



Step 4 Now move back to the original arbitrary point and use the distance formula, along with the results of steps 1, 2, and 3 to derive the equation of the ellipse

$$\frac{x^2}{a^2} + \frac{y^2}{b^2} = 1$$



At this point, I would also introduce the eccentricity so that we can work with the data that is given by astronomers.

$$\text{Eccentricity: } e = c / a$$

Now we are ready to put these equations to use. Transparencies 5 and 6 lead into the interesting fact that, even though the orbits of the planets are ellipses, they are very close to circles.

Group Project

Use the equations below to fill in the table.

$$\frac{x^2}{a^2} + \frac{y^2}{b^2} = 1 \quad a^2 = b^2 + c^2$$

$$e = \frac{c}{a}$$

Planet	Semimajor Axis a AU	Eccentricity e	Distance from Center to Focus c	Semiminor Axis b
Mercury	0.3870	0.2056		
Venus	0.7233	0.0068		
Earth	1.0000	0.0167		
Mars	1.5237	0.0934		
Jupiter	5.2034	0.0484		
Saturn	9.5371	0.0540		
Uranus	19.1913	0.0472		
Neptune	30.0690	0.0086		
Pluto	39.4817	0.2490		

Transparency 5

Group Project

Use the equations below to fill in the table.

$$\frac{x^2}{a^2} + \frac{y^2}{b^2} = 1 \quad a^2 = b^2 + c^2$$

$$e = \frac{c}{a}$$

Planet	Semimajor Axis a AU	Eccentricity e	Distance from Center to Focus c	Semiminor Axis b
Mercury	0.3870	0.2056	0.0796	0.3787
Venus	0.7233	0.0068	0.0049	0.7233
Earth	1.0000	0.0167	0.0167	0.9999
Mars	1.5237	0.0934	0.1423	1.5170
Jupiter	5.2034	0.0484	0.2518	5.1973
Saturn	9.5371	0.0540	0.5150	9.5232
Uranus	19.1913	0.0472	0.9058	19.1699
Neptune	30.0690	0.0086	0.2586	30.0679
Pluto	39.4817	0.2490	9.8309	38.2382

Transparency 6

The Orbit of UB 313

New 'planet' put to vote

The recent discovery of 2003 UB313 – an object larger and farther from the sun than Pluto – has scientists debating how to define the planet, if it should be one and what it might be named.

Orbit distance from the sun:
Closest: 39 AU
Farthest: 97 AU

NOTE: Orbic plane shows planet's position as of Aug. 10.

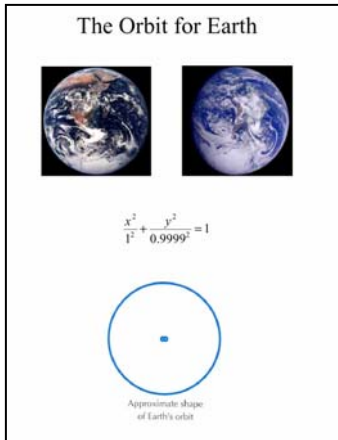
SOURCES: NASA; California Institute of Technology

Use the information on orbit distance from the sun to the equation of the orbit, and the eccentricity, for UB 313.

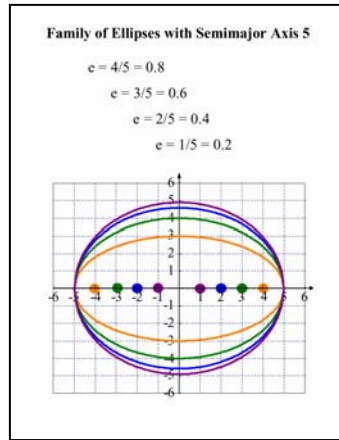
Transparency 20

Conic Sections (Continued)

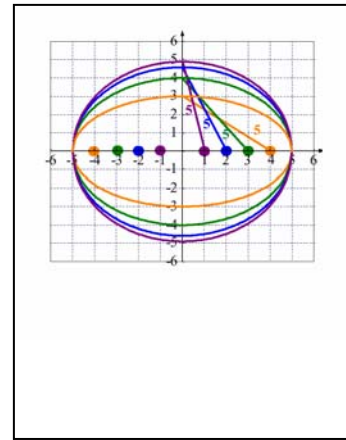
Transparency 7 shows two nice views of earth, along with the equation and the graph of its orbit. Transparencies 8 and 9 give us a chance to discuss eccentricity and reinforce the geometric relationship between a , b , and c .



Transparency 7

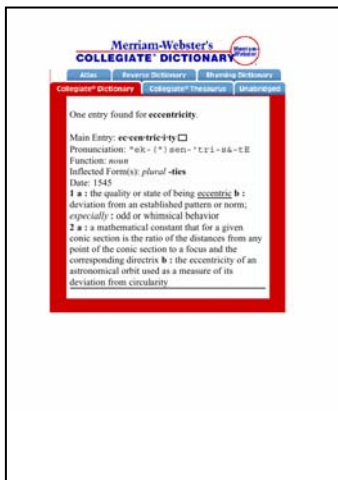


Transparency 8

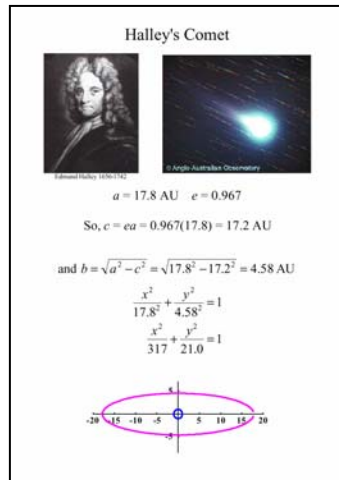


Transparency 9

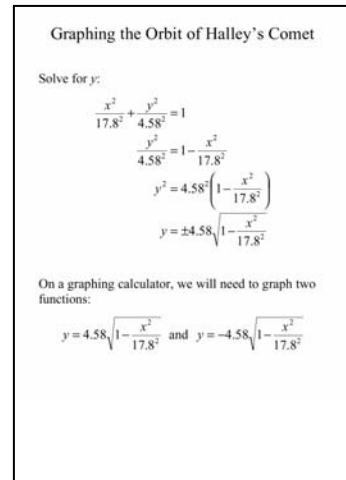
The definition for eccentricity in Transparency 10 is from the Internet (and the second part of the definition is what we use to introduce conic sections in polar coordinates). To show an orbit with an eccentricity close to 1, we can use Halley's Comet in Transparency 11. To graph the orbit of Halley's Comet on a graphing calculator we need to solve for y , as shown in Transparency 12.



Transparency 10



Transparency 11

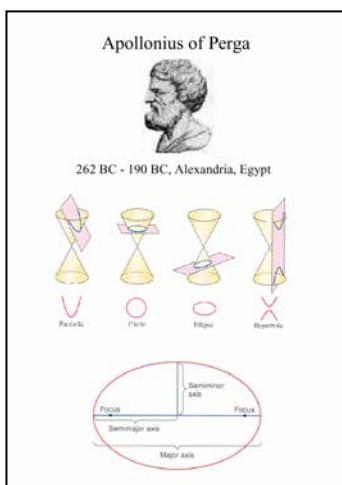


Transparency 12

As you know, it became apparent to me that I don't cover eccentricity in as much detail as I should in intermediate or college algebra. Now that I see how the astronomers give information about orbits, I am going to do more with it in my classes. These overheads may be a good way to start.

Conic Sections (Continued)

Now we are ready to introduce the rest of the conics. Transparency 13 does this very nicely. The idea is to show that conic sections have been studied many years before there were any applications discovered; the value of mathematics is not necessarily in its applications. It is interesting to note that, at the same time Apollonius was studying conic sections, the first humans were rowing their way to Hawaii.



Transparency 13

Arc Length

In calculus, after we have introduced the formulas for arc length, we can bring back Halley's Comet and use the graphing calculator to approximate the distance traveled by Halley's Comet in one trip around its orbit. This is shown in Transparency 14. Remember, we approximated the integral using the graphing calculator.

Find the distance traveled by Halley's Comet in one trip around its orbit

Arc Length Formula for $y = f(x)$

$$s = \int_a^b \sqrt{1 + (f'(x))^2} dx$$

Find the derivative:

$$y = 4.58 \sqrt{1 - \frac{x^2}{17.8^2}}$$
$$y' = 4.58 \frac{1}{2} \left(1 - \frac{x^2}{17.8^2}\right)^{-1/2} \left(\frac{-2x}{17.8^2}\right)$$
$$= \frac{-4.58x}{17.8^2} \left(1 - \frac{x^2}{17.8^2}\right)^{-1/2}$$

One-fourth of the distance will be

$$s = \int_0^{17.8} \sqrt{1 + \frac{4.58^2 x^2}{17.8^4} \left(1 - \frac{x^2}{17.8^2}\right)^{-1}} dx$$

Transparency 14

Conic Sections in Polar Coordinates

Use the online dictionary definition for eccentricity as a lead-in to polar equations for conics. The diagram in Transparency 15 is drawn from this definition. Transparency 16 shows an ellipse drawn from the online definition, with a coordinate system superimposed on it so that the origin of the coordinate system is at the focus. Point P has polar coordinates (r, θ) . Drop a line segment from P to the x -axis at a right angle. This forms a triangle with base $r \cos \theta$. Transparency 17 shows how we manipulate this formula to find an equation for the conic with r given in terms of θ .

Transparency 15

Transparency 15

We superimpose a coordinate system on our ellipse, with the origin at the focus.

Then we label point P with polar coordinates (r, θ) .

The definition, together with the labels on our diagram gives us:

$$e = \frac{|PF|}{|PD|} = \frac{r}{d - r \cos \theta}$$

Transparency 16

Transparency 16

Solve for r :

$$ed - r \cos \theta = r$$

$$ed - r \cos \theta = r$$

$$r + r \cos \theta = ed$$

$$r(1 + \cos \theta) = ed$$

$$r = \frac{ed}{1 + \cos \theta}$$

Transparency 17

Transparency 17

Transparency 18 can be used to derive the equation of the conic in terms of only a and e , the form used by astronomers. Transparency 19 shows how we did this during the talk.

Writing the equation in terms of e and a .

$$r = \frac{ed}{1 + e \cos \theta}$$

Transparency 18

Transparency 18

Writing the equation in terms of e and a .

$$r = \frac{ed}{1 + e \cos \theta}$$

When $\theta = 0$, $r_1 = \frac{ed}{1+e}$

When $\theta = \pi$, $r_2 = \frac{ed}{1-e}$

but $r_1 + r_2 = \frac{ed}{1+e} + \frac{ed}{1-e} = 2a$

$$ed(1-e) + ed(1+e) = 2a(1-e^2)$$

$$ed(1-e+1+e) = 2a(1-e^2)$$

$$2ed = 2a(1-e^2)$$

Therefore, $r = \frac{a(1-e^2)}{1+e \cos \theta}$

Transparency 19

Transparency 19

I hope you find these transparencies useful

Web Links for Conic Sections talk:

The first two links are to the demonstrations of Kepler's laws that were shown during the talk. They come from Dr. Les Tomley at San Jose State University.

Kepler's first and second laws:

<http://www.physics.sjsu.edu/tomley/Kepler12.html>

Kepler's third law:

<http://www.physics.sjsu.edu/tomley/Kepler3.html>

Solar System animation:

<http://astro.u-strasbg.fr/~koppen/orbitviewer/halley.html>

Postage Stamps of Mathematicians

<http://jeff560.tripod.com/>

Powers of 10 animation:

<http://micro.magnet.fsu.edu/primer/java/scienceopticsu/powersof10/index.html>